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High-performance, compact optical standard

Zachary L. Newman,^{1,*,†} Vincent Maurice,^{1,†} Connor Fredrick,^{1,2} ^(D) Tara Fortier,¹ Holly Leopardi,¹ Leo Hollberg,³ Scott A. Diddams,^{1,2} ^(D) John Kitching,¹ and Matthew T. Hummon¹ ^(D)

¹Time and Frequency Division, National Institute of Standards and Technology, 325 Broadway, Boulder, Colorado 80305, USA ²Department of Physics, University of Colorado, 2000 Colorado Ave., Boulder, Colorado 80309, USA ³Department of Physics, Stanford University, 382 Via Pueblo Mall, Stanford, California 94305-4060, USA *Corresponding author: zachary.newman@octavephotonics.com

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We describe a high-performance, compact optical frequency standard based on a microfabricated Rb vapor cell and a low-noise, external cavity diode laser operating on the Rb two-photon transition at 778 nm. The optical standard achieves an instability of $1.8 \times 10^{-13} \tau^{-1/2}$ for times less than 100 s and a flicker noise floor of 1×10^{-14} out to 6000 s. At long integration times, the instability is limited by variations in optical probe power and the ac Stark shift. The retrace was measured to 5.7×10^{-13} after 30 h of dormancy. Such a simple, yet high-performance optical standard could be suitable as an accurate realization of the meter or, if coupled with an optical frequency comb, as a compact atomic clock comparable to a hydrogen maser. © 2021 Optical Society of America

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Two-photon transitions in alkali and alkaline Earth metals are considered promising candidates for simple, field-deployable optical frequency references and optical clocks [1,2]. To first order, Doppler-free spectroscopy can be accomplished by driving the transition using counterpropagating laser beams at the same frequency, resulting in spectrally narrow fluorescence signals. Because atoms in all velocity classes contribute to the signal and because two-photon transitions are typically observed in fluorescence through decay channels at a wavelength distinct from the driving field, measurements of such transitions can be performed with high signal-to-noise ratios. As a result, optical standards based on two-photon transitions can achieve excellent stability and accuracy.

The two-photon transition in Rb at 778 nm was studied extensively as an optical frequency reference throughout the 1990s and early 2000s [3–5]. The short-term frequency stability of these early tabletop systems [4,5] was measured at $\approx 3 \times 10^{-13} \tau^{-1/2}$ out to 1000 s. More recently, improved longterm performance has been demonstrated at $\approx 4 \times 10^{-13} \tau^{-1/2}$ out to 10,000 s [6]. Dichroic standards are also being developed, in which much smaller detunings from the intermediate level result in larger signals, at the cost of increased light shift [7,8]. With the recent progress in miniaturization of frequency comb technology (both fiber and microresonator based) there has been renewed interest in the Rb two-photon transition for use as the local oscillator in a compact low-power optical clock [6,9]. To that end, our group recently demonstrated a miniature optical reference, with the optical bench having a total volume of 35 cm³, and performance of $3 \times 10^{-12} \tau^{-1/2}$, limited by the frequency noise of the semiconductor probe laser [10].

In this Letter, we stabilize a near-infrared, external cavity diode laser (ECDL) to the two-photon transition in rubidium-85 confined in a microfabricated vapor cell. We achieve an instability of $1.8 \times 10^{-13} \tau^{-1/2}$ from 1 s to 100 s, comparable to or outperforming previous measurements of the two-photon standard over these time scales. In the current configuration, the laser achieves an instability of $\approx 1 \times 10^{-14}$ at 1000 s.

Figure 1 shows a schematic of the optical standard, which consists of an ECDL stabilized to the $5S_{1/2}(F=3) \rightarrow$ $5D_{5/2}(F=5)$, two-photon transition in rubidium-85. Figure 1(b) gives a close-up view of the vapor cell, which employs an in-line geometry for probing and measuring the two-photon transition [9], in which the counterpropagating beams necessary for avoiding Doppler broadening of the transition are generated by retro-reflecting the laser off a high-reflectivity dielectric coating on the back of a planar, microfabricated cell. We introduce rubidium (natural abundance) into the cell by laser activation of a rubidium dispenser compound after the cell is sealed. While this introduces some unwanted background gases, it serves as a convenient way to load atoms into a mm-scale cell. The addition of a nonevaporable getter helps preserve good vacuum inside the chamber that allows us to do high-resolution spectroscopy and helps mitigate some of the shifts associated with background gases from the dispenser.

A photomultiplier tube (PMT) placed at the back of the cell detects excitation of the transition via fluorescence at 420 nm. In contrast, laboratory-scale rubidium two-photon standards typically locate the PMT adjacent to the vapor cell [5,6,11]. The compact cell geometry also reduces the sensitivity of average probe beam power due to angular misalignment of the beam steering optics, as the sensitivity scales as roughly the square of the beam propagation length for the simple retro-reflection geometry [12].



Fig. 1. (a) Optical standard consists of a 778 nm ECDL, power-stabilized using an acousto-optical modulator (AOM), a microfabricated vapor cell housed in a magnetic shield, and a photomultiplier tube (PMT). The laser is locked to the two-photon transition in Rb at 778 nm using analog electronics, and its frequency is measured with respect to a titanium-doped sapphire (henceforth, Ti:sapphire) frequency comb stabilized to an optical cavity and hydrogen maser at short and long integration times, respectively. (b) Microfabricated vapor cell has optically coated windows, and an aspheric collection optic directs the blue fluorescence from the atoms onto the PMT. (c) Spectrum of two of the hyperfine components of the two-photon transition (blue) and the corresponding fit (orange) showing a linewidth of ≈ 680 kHz FWHM.

We measure the frequency of the ECDL by comparing the heterodyne beat between the ECDL and a mode of a selfreferenced, 1 GHz repetition-rate, titanium sapphire frequency comb [13] locked to an ultrastable optical cavity [14]. At short integration times, the frequency comb stability, determined by the stability of the reference cavity ($\sigma \approx 1 \times 10^{-15}$ at 1 s), is more than sufficient for measuring the frequency of the ECDL. However, at long integration times, linear drift of the cavity is a limiting factor in our measurements. To work around this issue, we track the cavity drift by measuring the repetition of the Ti:sapphire laser against a maser-synthesized 1 GHz tone and subtract a fourth-order polynomial fit of the frequency drift from the ECDL-comb beat in the optical domain.

Figure 1(c) shows a spectrum (blue dots) of the hyperfine components (F = 5 to 1) of the clock transition as the laser frequency is swept across the resonance with a fit (orange line) that gives a linewidth of 680 kHz. The observed linewidth includes contributions from the natural linewidth (330 kHz), transit time broadening (\approx 110 kHz) resulting from the finite interaction time of the atoms with the laser [15], (\approx 10 kHz) due to self-collisional broadening [16], and (39 kHz) from the laser linewidth. The remaining \approx 190 kHz of broadening places a limit of unwanted background gas pressure in the cell of \approx 9 mTorr [16]. These contributions are consistent with earlier measurements of the same vapor cell [9].

Figure 2(a) shows a 2.5 h time series of the locked laser frequency from which we calculate the Allan deviation of the



Fig. 2. Short-term frequency stability of the optical standard. (a) Time-series measurement of the beat frequency between the Ti:sapphire frequency comb and the ECDL. (b) Allan deviations of the rubidium-stabilized clock laser frequency with 1σ error bars for a 16 h period (orange), the 2.5 h selection (blue) shown in (a). (c) Frequency noise power spectrum of the free-running, unmodulated ECDL. The shaded area indicates the component of the spectrum that contributes dominantly to the linewidth [17]. The vertical red line indicates twice the modulation frequency. (d) Scaling of the 1 s instability with input power. Star denotes operating power for (a) and (b). The dashed line indicates a fit of the instability limit due to shot noise and intermodulation noise.

optical standard, shown in Fig. 2(b). The typical performance of the system is well characterized by the 16 h data set (orange). At longer times (≈ 1000 s), we observe a flicker noise floor of $\approx 1 \times 10^{-14}$, and attribute this to ac Stark shifts caused by instability of the laser probe intensity as described in the following section. During time periods where the intensity of the driving field is more stable, however, we commonly see improved performance as exemplified by the 2.5 h data set (blue).

We first consider the short-term instability, which scales as $1.8 \times 10^{-13} \tau^{-1/2}$ for times less than 100 s. There are two main contributions to this instability that we assess in (c) and (d); (c) shows the measured frequency noise spectrum $S_{\Delta\nu}(f)$ of the free-running ECDL. According to [18], the short-term instability limit imposed by frequency noise of the local oscillator (intermodulation noise) is $\sigma_y \approx \frac{[S_{\Delta\nu}(2f_m)/f_0^2]^{1/2}}{2} \frac{1}{\sqrt{\tau}}$, where f_m corresponds to the dither frequency used to generate the error signal via lock-in detection, and f_0 is the clock laser frequency. For our system, this corresponds to $S_{\Delta\nu} = 5 \times 10^3 \text{ Hz}^2/\text{Hz}$ and an instability of $9 \times 10^{-14} \tau^{-1/2}$. For reference, we modulate the ECDL frequency directly via the laser current, and integration of the free-running frequency noise of the laser below the β -line (shaded red region) yields a 39 kHz linewidth [17].

Figure 2(d) shows the scaling of the short-term instability at 1 s averaging time with input probe power P_{in} . For an instability



Fig. 3. Systematic contributions to the instability of the optical standard. (a) Frequency shift versus input power. (b) Frequency shift versus cell temperature. (c) Allan deviations showing the measured frequency data (orange, from Fig. 2) and the expected contribution to the clock frequency from the light shift (purple, dark blue) and the cell temperature shift ("tempco," green). Error bars represent a 1σ confidence interval.

limited by fluorescence photon shot noise, we expect the instability to scale as P_{in}^{-1} , which agrees well with the observations for powers less than 10 mW. The dashed line in Fig. 2(d) is a fit to the instability data where a term that scales as P_{in}^{-1} and a constant offset of 9×10^{-14} , corresponding to the intermodulation limit, have been added in quadrature. This model fits the instability data well and indicates that operating at powers larger than 10 mW does not further improve the short-term instability due to the limit imposed by intermodulation noise for this laser. The time-series data presented in Fig. 2 are taken with an input power of ≈ 6 mW, as denoted by the star in 2(d).

Figure 3 summarizes measurements of systematic effects that limit the stability of the optical reference at long times. Figures 3(a) and 3(b) show measurements of the ac Stark shift and vapor cell temperature shift, respectively. In 3(a) and 3(b), the input power and cell temperature are recorded using in-loop control measurements. The observed shifts are consistent with previous measurements [6]. Figure 3(c) shows the limits to the laser instability imposed by the ac Stark shift and cell temperature variation as recorded by out-of-loop measurements for the laser power and cell temperature [PD2 as shown in Fig. 1(b) and an auxiliary thermistor placed in the cell mount].

A measurement of the 420 nm fluorescence amplitude, which is proportional to the square of the probe power, provides a second out-of-loop monitor of the laser power. Figure 3(c)shows fractional instabilities for both out-of-loop measurement techniques. We find that both out-of-loop power measurements overestimate the actual variation observed in operation of the optical standard at integration times longer than 1000 s.



Fig. 4. History of measurements of the ⁸⁵Rb two-photon transition. The purple star (dark gray dashed line) indicates the BIPM (2005) accepted value for the $F = 3 \rightarrow F' = 5$ transition. The number above each measurement indicates the vapor cell temperature in °C. Error bars for this work represent uncertainty due to operating temperature range. For previous measurements, refer to original papers for description of error bars.

Figure 4 shows prior measurements of ⁸⁵Rb, $F = 3 \rightarrow F' = 5$ transition corrected for zero optical power [3,19–24]. Taking the mean value of our 16 h measurement campaign, and correcting for zero optical power and 95°C cell temperature, we measure the frequency of our optical standard to be \approx (385 285 142 369 \pm 5) kHz (red circle). This corresponds to a frequency difference $\Delta v \approx -6 \pm 5$ kHz from the Bureau International des Poids et Mesures (BIPM) (2005) accepted transition frequency [20,21]. The uncertainty of \pm 5 kHz corresponds to the cell temperature shift over the typical operating temperature range of 90°C to 100°C.

Frequency shifts due to collisions with background gas impurities typically limit the accuracy of vapor cell optical references [4,23]. Previous measurements made using fused silica or pyrex cells are likely to contain 4 mTorr of helium due to permeation of atmospheric helium into the cell, leading to a frequency shift of +4 kHz compared to an evacuated cell [16]. In contrast, the vapor cell used in this work is fabricated using low-heliumpermeability aluminosilicate glass windows, with a predicted helium leak rate [25] of 1.2×10^{-5} Torr/year, and negligible frequency shift due to collisions with helium. Our rubidium cell loading method introduces additional background gases into the cell that we subsequently attempt to remove using pumping from a non-evaporable getter. Thus, estimating the additional background gas impurities after the cell is sealed is difficult, but our two-photon frequency measurement is consistent with a pressure of less than 2 mTorr.

Figure 5 shows a 24 h retrace measurement. Retrace, as described by Vanier et al. [26], in the context of our optical standard refers to a change in the clock laser frequency after the system power has been cycled and is particularly relevant for battery-powered or portable frequency standards. Figure 5 shows a fractional retrace measurement for the optical standard with a 5-24-5-5 h on-off -warm-up-on cycle. For this measurement, the optical standard was operated at 0.7 mW probe power. After a warm-up period of 5 h, we achieve a retrace of 5.7×10^{-13} , which represents an improvement of close to 10^3 compared to the retrace of the chip-scale atomic clock [27]. The warm-up time of 5 h is currently limited by thermalization of the optical bread board base plate, which could be reduced in the future by implementing a more compact physics package. This improvement in retrace will be important for applications where a portable clock needs to achieve a high level of accuracy after turn on when disciplining to a reference oscillator (e.g., global navigation satellite system timing) is impracticable.

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Fig. 5. Fractional retrace measurement showing the (a) clock laser frequency and (b) cell housing temperature over a $\approx 2 d$ period. After 10 h initial measurement, the laser and all control electronics are powered off, left in the "off" state for over a day, and powered on again. (a) Beat note between the ECDL and an auxiliary erbium fiber comb. Highlighted segments of the data indicate the portion of the data used to determine the retrace, allowing for a ≈ 4 h warm-up period over which environmental factors surrounding the cell equilibrate (the cell housing temperature is shown as an example). Dashed-dotted lines show the average frequency of the blue and orange measurement highlighted segments. (c) Fractional stability of the laser frequencies shown in (a). Fractional instability at 1000 s, $\approx 1 \times 10^{-13}$, corresponds to the uncertainty in our retrace measurements.

In summary, we have demonstrated a high-quality optical standard and presented a framework for future evaluations of compact optical references (absolute frequency combined with retrace measurements). Based on our measurements, we believe optical standards based on optical transitions in warm alkali vapors, such as rubidium, can support even higher performance; in fact, we have observed 8×10^{-14} at 1 s using a low-noise clock laser. We suspect short-term instabilities at the 1×10^{-14} level are within reach using a high-power, low-noise laser in conjunction with a light-shift mitigating interrogation scheme [28].

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Data Availability. Data underlying the results presented in this paper are not publicly available at this time but may be obtained from the authors upon reasonable request.

REFERENCES

1. G. Grynberg and B. Cagnac, Rep. Prog. Phys. 40, 791 (1977).

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- 2. J. L. Hall, M. Zhu, and P. Buch, J. Opt. Soc. Am. B 6, 2194 (1989).
- F. Nez, F. Biraben, R. Felder, and Y. Millerioux, Opt. Commun. 102, 432 (1993).

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- L. Hilico, R. Felder, D. Touahri, O. Acef, A. Clairon, and F. Biraben, Eur. Phys. J. Appl. Phys. 4, 219 (1998).
- M. Poulin, C. Latrasse, D. Touahri, and M. Têtu, Opt. Commun. 207, 233 (2002).
- K. W. Martin, G. Phelps, N. D. Lemke, M. S. Bigelow, B. Stuhl, M. Wojcik, M. Holt, I. Coddington, M. W. Bishop, and J. H. Burke, Phys. Rev. Appl. 9, 014019 (2018).
- C. Perrella, P. S. Light, J. D. Anstie, F. N. Baynes, R. T. White, and A. N. Luiten, Phys. Rev. Appl. 12, 54063 (2019).
- 8. V. Gerginov and K. Beloy, Phys. Rev. Appl. 10, 014031 (2018).
- 9. Z. L. Newman, V. Maurice, and T. E. Drake, et al., Optica 6, 680 (2019).
- V. Maurice, Z. L. Newman, S. Dickerson, M. Rivers, J. Hsiao, P. Greene, M. Mescher, J. Kitching, M. T. Hummon, and C. Johnson, Opt. Express 28, 24708 (2020).
- J. E. Bernard, A. A. Madej, K. J. Siemsen, L. Marmet, C. Latrasse, D. Touahri, M. Poulin, M. Allard, and M. Têtu, Opt. Commun. **173**, 357 (2000).
- K. W. Martin, B. Stuhl, J. Eugenio, M. S. Safronova, G. Phelps, J. H. Burke, and N. D. Lemke, Phys. Rev. A 100, 023417 (2019).
- 13. T. M. Fortier, A. Bartels, and S. A. Diddams, Opt. Lett. **31**, 1011 (2006).
- B. C. Young, F. C. Cruz, W. M. Itano, and J. C. Bergquist, Phys. Rev. Lett. 82, 3799 (1999).
- 15. F. Biraben, M. Bassini, and B. Cagnac, J. Phys. 40, 445 (1979).
- N. D. Zameroski, G. D. Hager, C. J. Erickson, and J. H. Burke, J. Phys. B 47, 225205 (2014).
- 17. G. Di Domenico, S. Schilt, and P. Thomann, Appl. Opt. 49, 4801 (2010).
- C. Audoin, V. Candelier, and N. Dimarcq, IEEE Trans. Instrum. Meas. 40, 121 (1991).
- R. Felder, D. Touahri, O. Acef, L. Hilico, J.-J. Zondy, A. Clairon, B. de Beauvoir, F. Biraben, L. Julien, F. Nez, and Y. P. Millerioux, Proc. SPIE 2378, 52 (1995).
- D. Touahri, O. Acef, A. Clairon, J.-J. Zondy, R. Felder, L. Hilico, B. de Beauvoir, F. Biraben, and F. Nez, Opt. Commun. 133, 471 (1997).
- D. J. Jones, S. A. Diddams, J. K. Ranka, A. Stentz, R. S. Windeler, J. L. Hall, and S. T. Cundiff, Science 288, 635 (2000).
- S. Diddams, D. Jones, J. Ye, S. Cundiff, J. Hall, J. Ranka, R. Windeler, R. Holzwarth, T. Udem, and T. Hänsch, Phys. Rev. Lett. 84, 5102 (2000).
- C. S. Edwards, G. P. Barwood, H. S. Margolis, P. Gill, and W. R. C. Rowley, Metrologia 42, 464 (2005).
- 24. O. Terra and H. Hussein, Appl. Phys. B 122, 27 (2016).
- A. T. Dellis, V. Shah, E. A. Donley, S. Knappe, J. Kitching, A. R. T. D. Ellis, V. I. S. Hah, E. L. A. D. Onley, S. V. K. Nappe, A. T. Dellis, V. Shah, E. A. Donley, S. Knappe, and J. Kitching, Opt. Lett. 41, 2775 (2016).
- J. Vanier, J. Gagnepain, W. Riley, F. Walls, and M. Granveaud, in IEEE Frequency Control Symposium (2003), pp. 807–815.
- "Microsemi space chip scale atomic clock data sheet," 2021, https:// www.microsemi.com/document-portal/doc_download/1243238space-csac-datasheet.
- V. Yudin, M. Y. Basalaev, A. Taichenachev, J. Pollock, Z. Newman, M. Shuker, A. Hansen, M. Hummon, R. Boudot, E. Donley, and J. Kitching, Phys. Rev. Appl. 14, 024001 (2020).

[†]These authors contributed equally to this Letter.